



# Wide-Band VLBA Receiver System for Astrometry and Navigation

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*US VLBA Navigation and Reference Frame Workshop*

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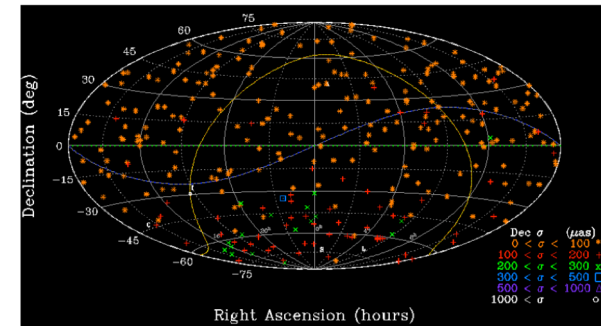
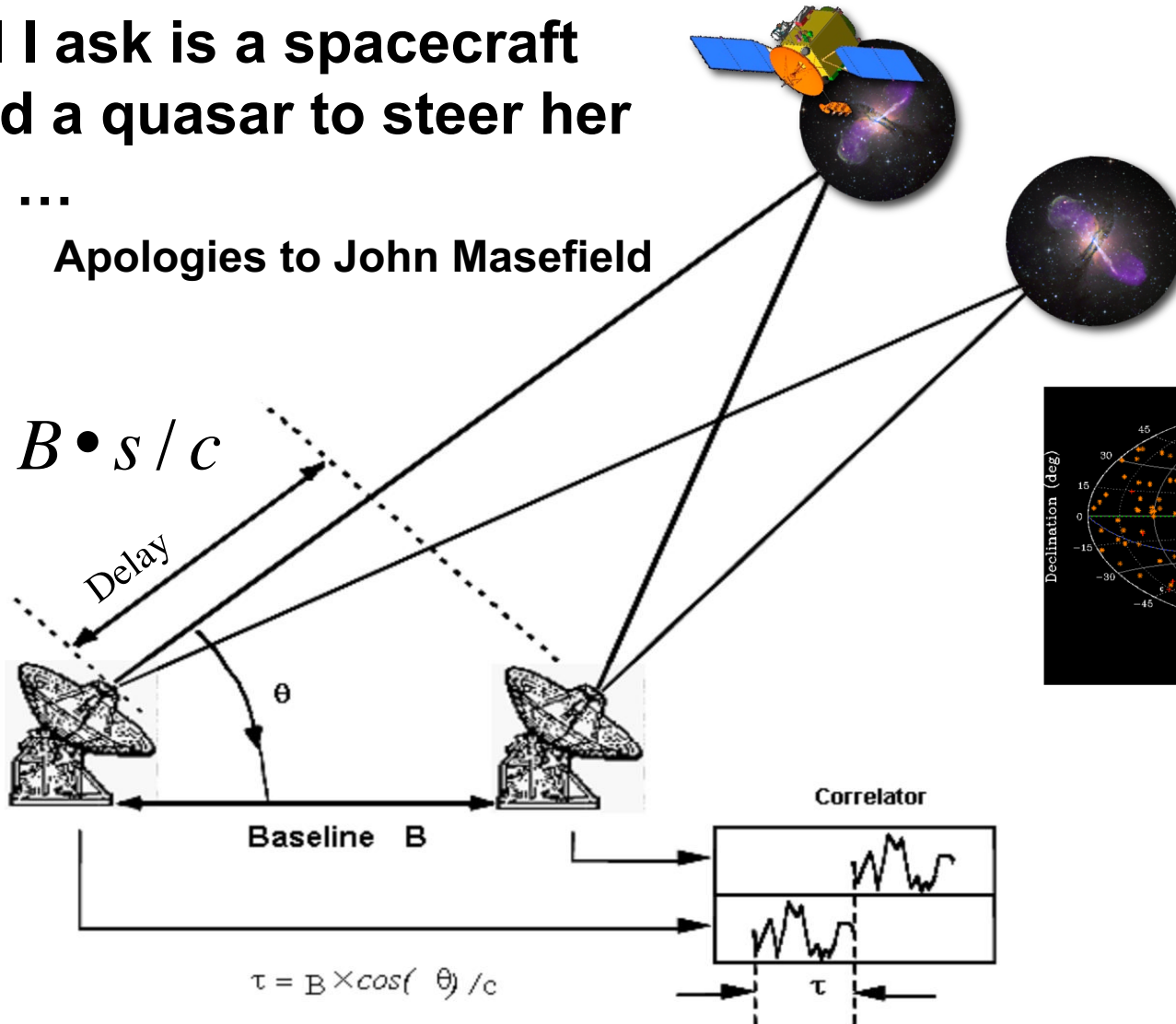
# Motivation I

Wide-Band VLBA Receiver System for Astrometry and Navigation

**All I ask is a spacecraft  
and a quasar to steer her  
by ...**

Apologies to John Masefield

$$\tau = B \cdot s / c$$



Credit: C.S. Jacobs



# Motivation I

## Wide-Band VLBA Receiver System for Astrometry and Navigation



# Motivation II

## Wide-Band VLBA Receiver System for Astrometry and Navigation

### Frequency Bands Allocated by the International Telecommunication Union (ITU)

Band	> 2 million km from Earth		< 2 million km from Earth	
	Uplink	Downlink	Uplink	Downlink
S	2110–2120*	2290–2300	2025–2110	2200–2290
X	7145–7190	8400–8450	7190–7235	8450–8500
K	-	-	-	25500–27000
Ka	34200–34700	31800–32300	-	-

DSN Telecommunications Link Design Handbook, 810-005, 201  
“Frequency and Channel Assignments”

**\* Deep Space S-band not available at Madrid Deep Space Communications Complex**



# VLBA vs. Spacecraft TT&C

## Telecommunications, Tracking, & Command

Table 3: Receiver Frequency Ranges & Performance  
as of July 2015

[1] Receiver Band Designation (*)	[2] Nominal Frequency Range [GHz]	[3] Typical Zenith SEFD [Jy]	[4] Center Frequency for SEFD [GHz]	[5] Typical Peak Gain [K Jy <sup>-1</sup> ]	[6] Baseline Sensitivity $\Delta S^{2048,2m}$ [mJy]	[7] Image Sensitivity $\Delta I_m^{2048,8h}$ [μJy beam <sup>-1</sup> ]
90 cm (a)	0.312 - 0.342	2742	0.326	0.077 (h)	39	(j) 266
50 cm (a,b)	0.596 - 0.626	2744	0.611	0.078 (h)	111	(k) 681
21 cm (c)	1.35 - 1.75	289	1.438	0.110	1.0	10
18 cm (c)	1.35 - 1.75	314	1.658	0.112	1.1	11
13 cm	2.2 - 2.4	347	2.269	0.087	1.2	12
13 cm (d)	2.2 - 2.4	359	2.269	0.085	1.3	12
6 cm (e)	3.9 - 7.9	210	4.993	0.119	0.7	6
7 ghz (e)	3.9 - 7.9	278	6.660	0.103	1.0	9
4 cm	8.0 - 8.8	327	8.419	0.118	1.2	11
4 cm (d)	8.0 - 8.8	439	8.419	0.105	1.6	15
2 cm	12.0 - 15.4	543	15.363	0.111	1.9	18
1 cm (f)	21.7 - 24.1	640	22.236	0.110	2.3	22
24 ghz (f)	21.7 - 24.1	534	23.801	0.118	1.9	18
7 mm	41.0 - 45.0	1181	43.124	0.090 (h)	6	40
3 mm (g)	80.0 - 90.0	4236	86.2	0.033 (i)	30	(l) 254

- No coverage of K- or Ka-band TT&C allocations
- ~ 35%  $T_{sys}$  degradation at X band w/ dichroic

# Objective

## Wide-Band VLBA Receiver System for Astrometry and Navigation

- **International Celestial Reference Frame is now multi-wavelength**
- **VLBA is key asset in constructing current and future instances of ICRF,**
- **... but no coverage at bands relevant to spacecraft tracking**
- **Beginning technology development toward prototype wide-band receiver system for reference frame development and maintenance, with eye toward testing on VLBA**



# Technology Development for the North America Array

Background I (FY16--FY18)

## North America Array Receiver Package

### System Requirements

- Continuous coverage for 8-48 GHz
- $T_{rx} = 30\text{--}40\text{ K}$   
 $T_{sys}$  of 34 K @ 10 GHz and 45 K @ 30 GHz required
- Dual polarization

### Assumption

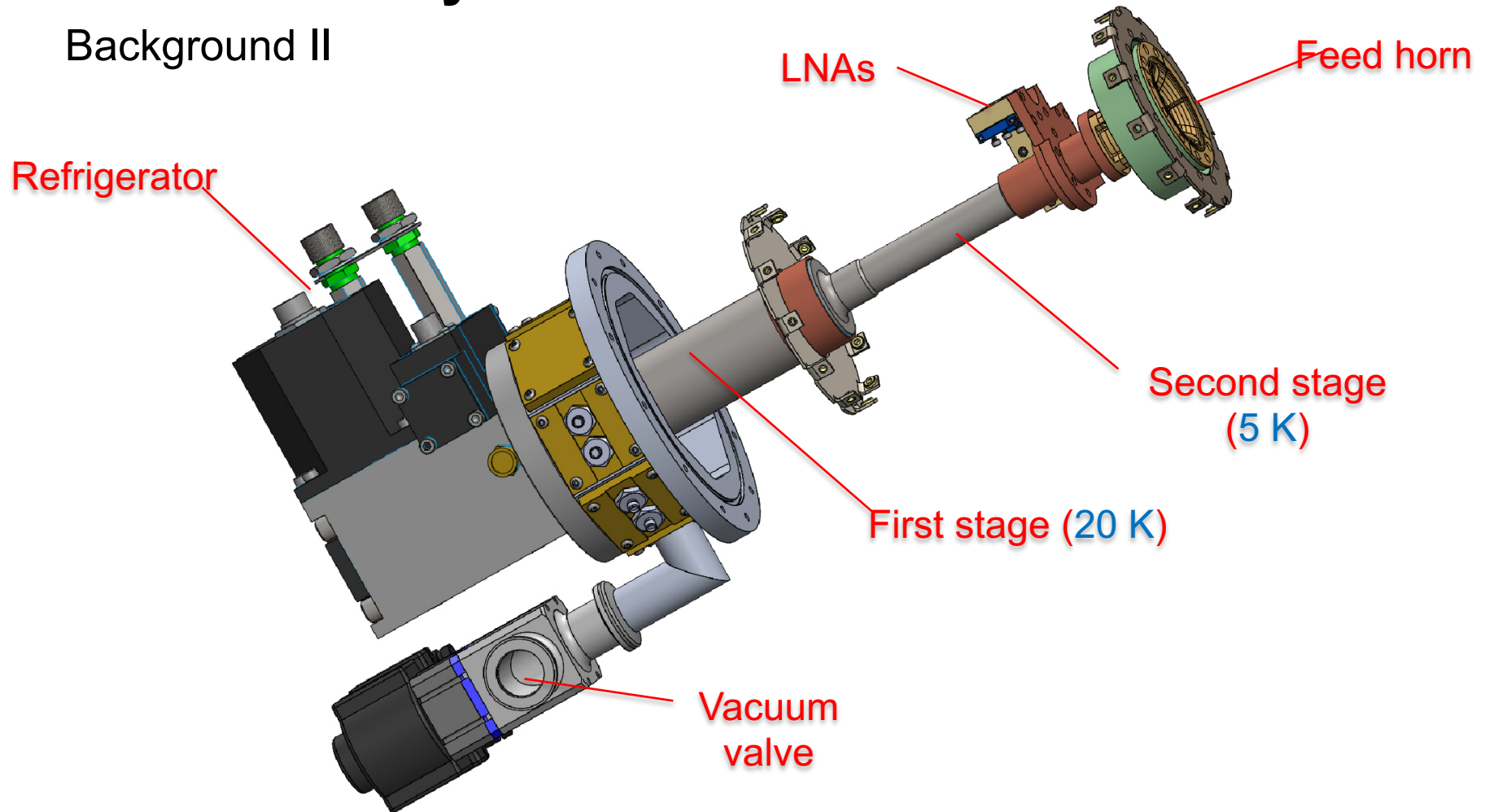
Baseline antenna is offset Gregorian (e.g., MeerKAT antenna) scaled to 18 m diameter with  $f/D = 0.55$

### Design considerations

- Compact cryogenic package
- Easy to manufacture
- Easy to service
- Low Cost

# Technology Development for the North America Array

## Background II

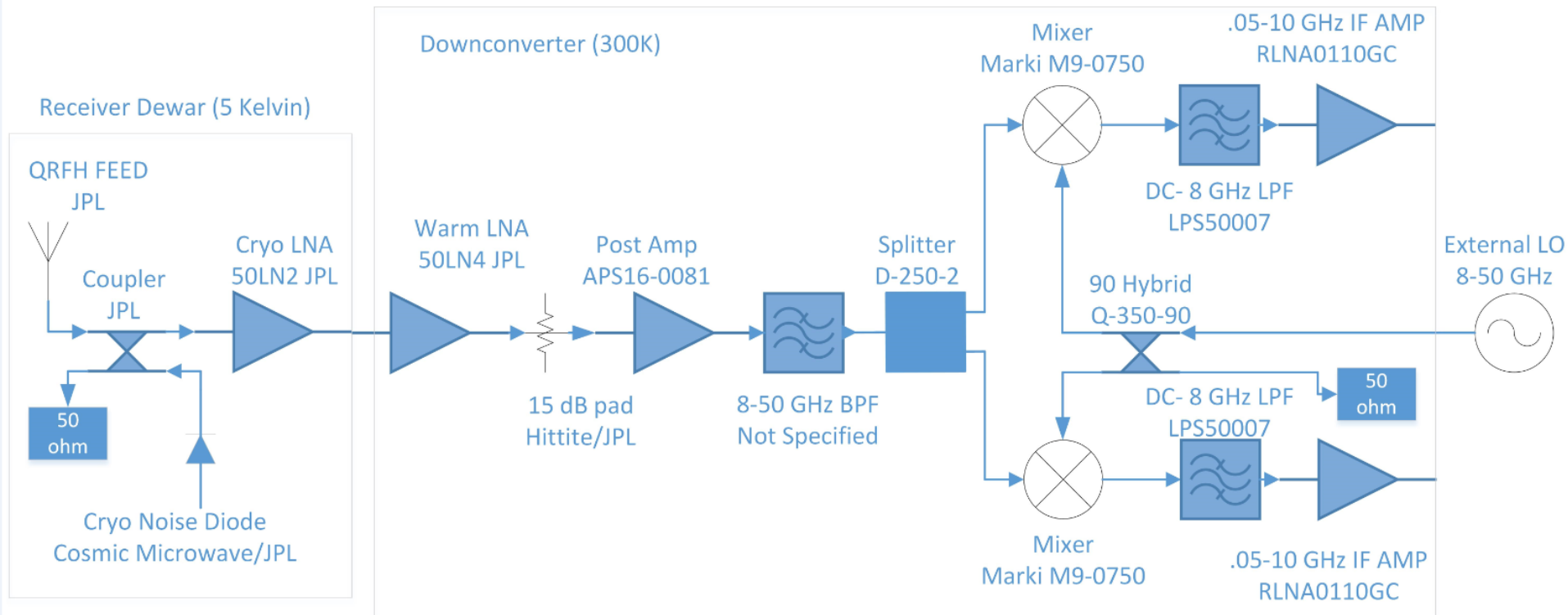




# Technology Development for the North America Array

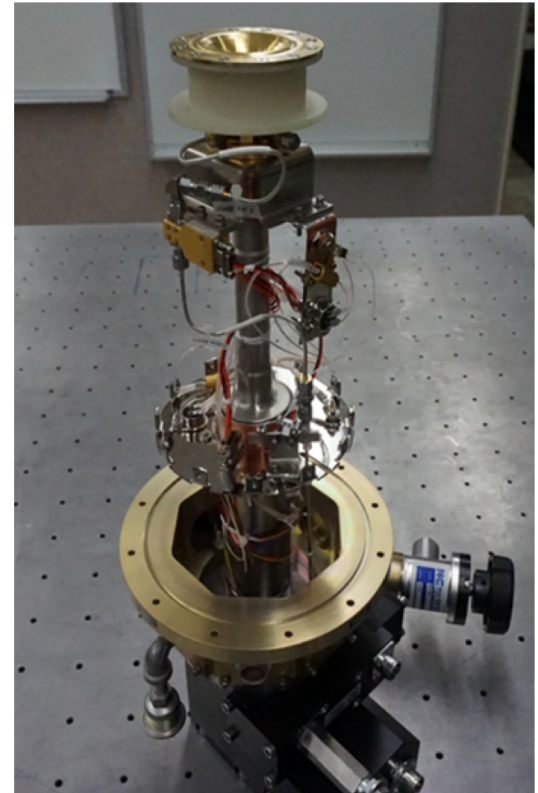
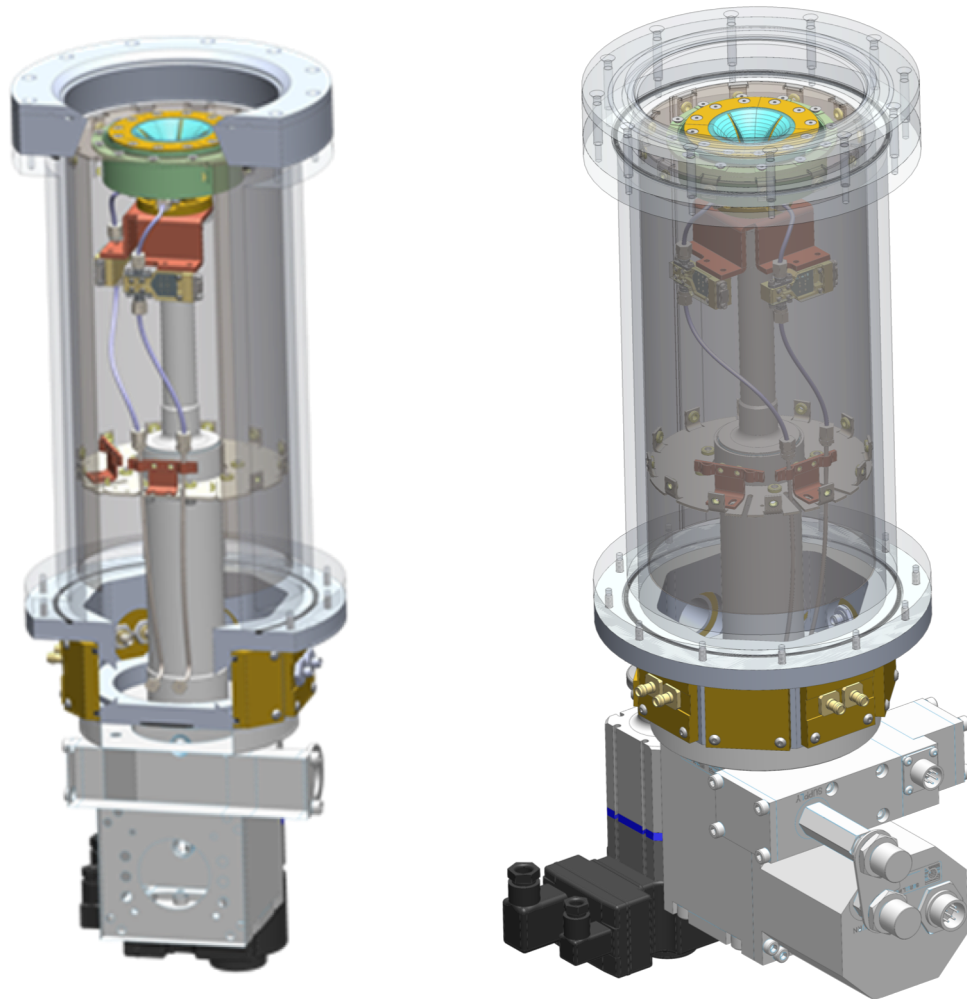
## Background III

### 8-48 GHz ultra wide band downconverter for the NAA



# Technology Development for the North America Array

## Background IV





# Technology Development for the North America Array

## Background V



# Wide-Band VLBA Receiver System for Astrometry and Navigation

## Requirements

### North America Array/next-generation Very Large Array

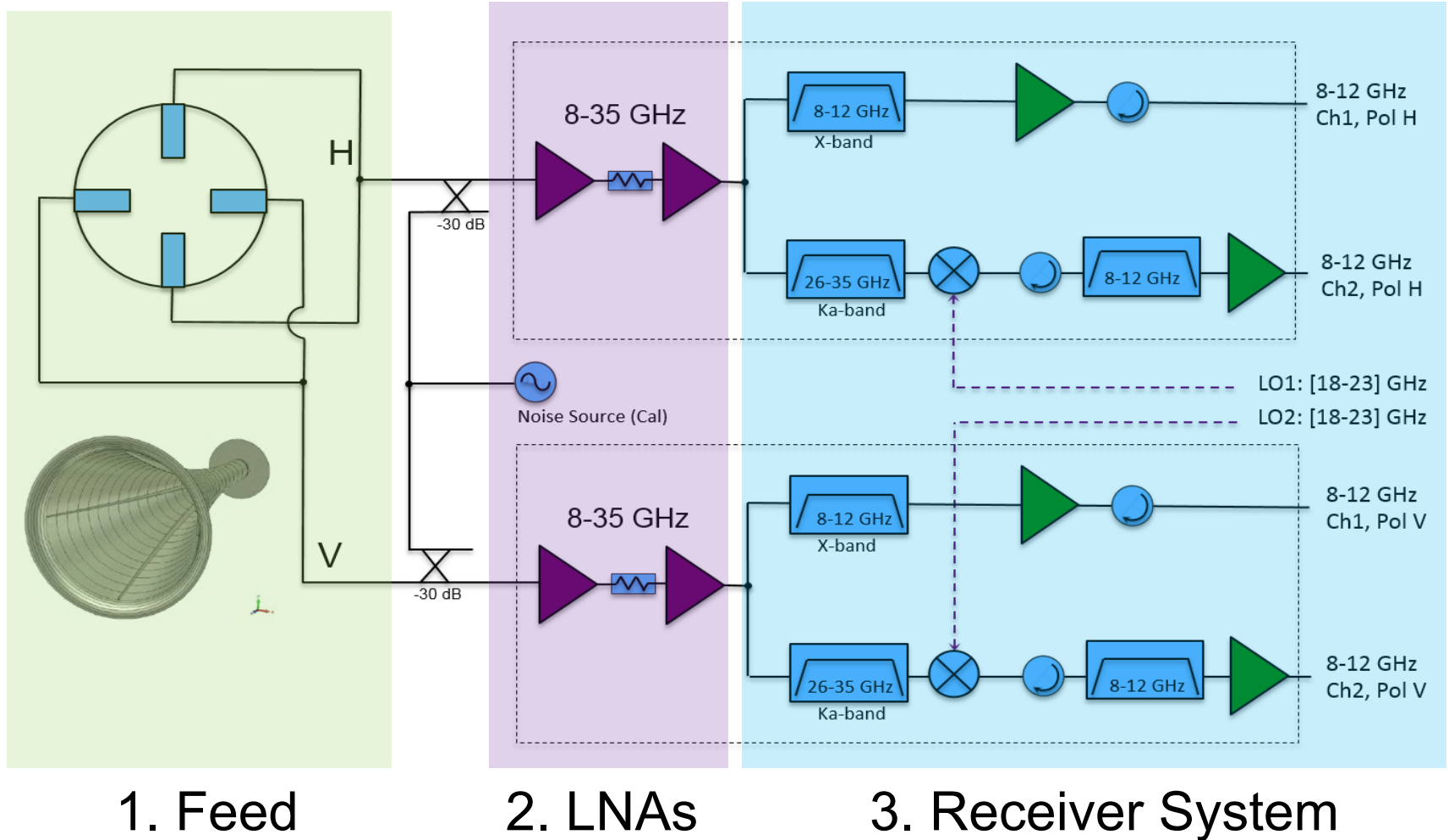
- Continuous coverage for 8-48 GHz
  - Science-driven requirement
- Baseline antenna is 18 m diameter offset Gregorian with  $f/D = 0.55$

### Very Long Baseline Array

- Coverage of 8 GHz to 35 GHz
  - Astrometry and navigation-driven requirement
  - Radar?
- Antenna is 25 m diameter Cassegrain

# VLBA Receiver System

## Three Technology Development Components

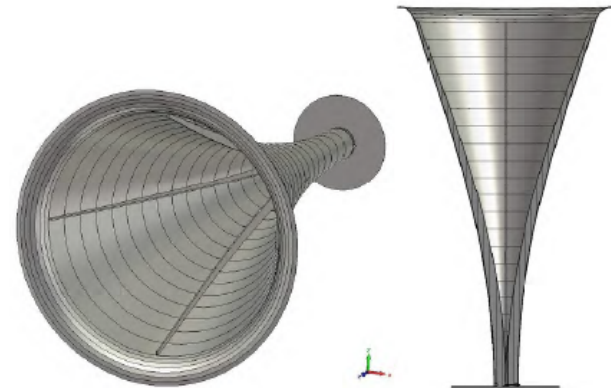


# Wideband Feed Development

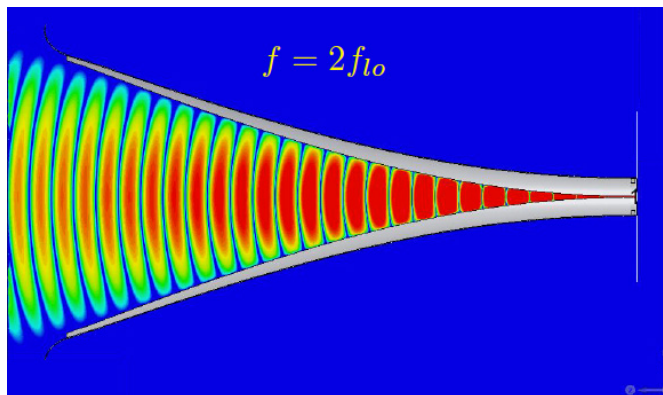
## Objectives and Challenges

- Design feed for 8-35 GHz
- Dual polarization
- Compatible with VLBA 25 m Cassegrain optics
- Quad-ridge flared horn (QRFH)?

Based on work of A. Akgiray & S. Weinreb



$\text{Gain} \sim 20 \text{ dB}$
$\overline{\eta}_{\text{apt}} \sim 65\%$
$X_{\text{pol}} 10 - 20 \text{ dB}$
$\text{IRL} > -10 \text{ dB}$



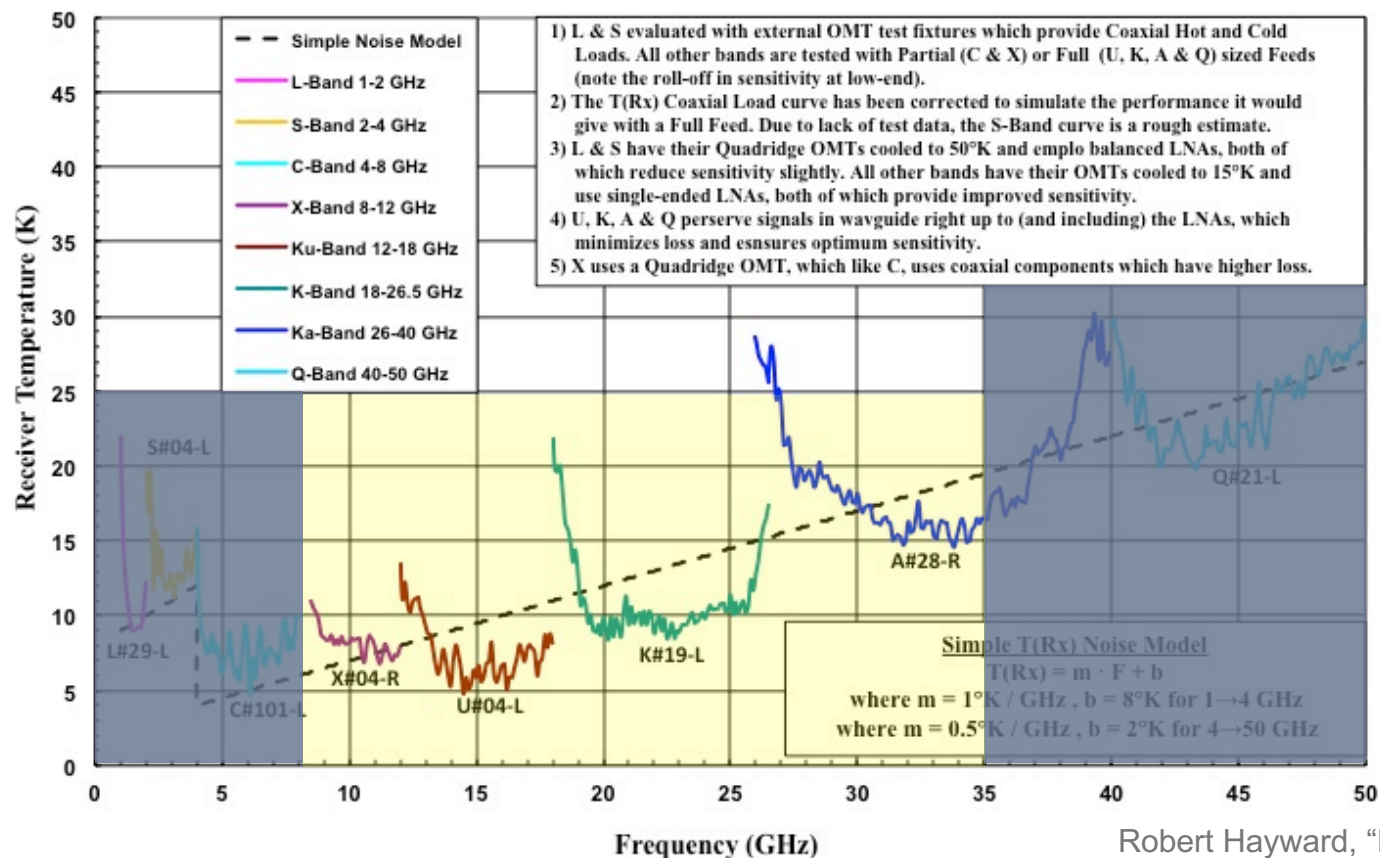
*Akgiray, PhD Thesis,  
Caltech 2013  
design intended for DSS-14  
with similar optics and f/D*



# LNA Development

Goal: Receiver Noise Comparable to JVLA

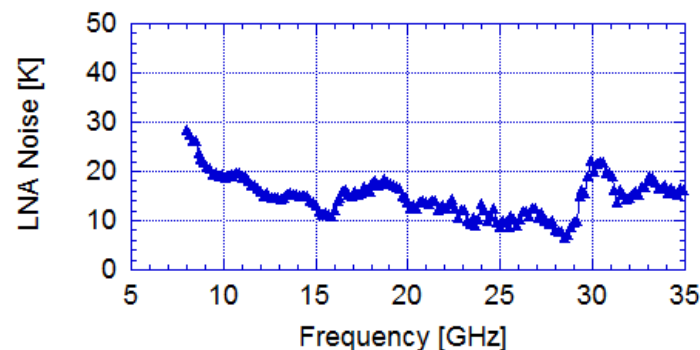
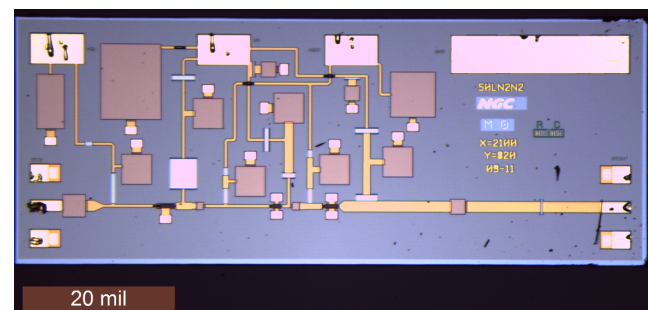
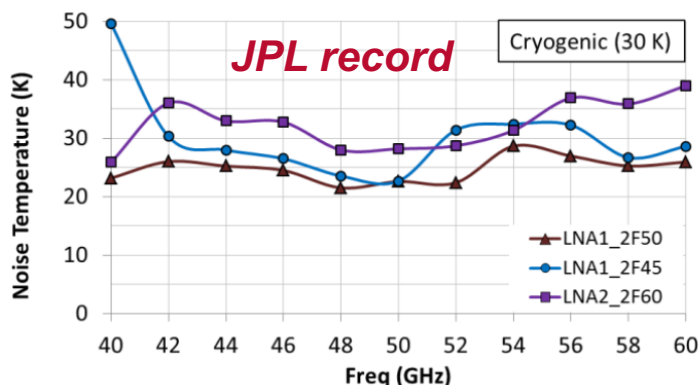
- HEMT LNA-based receivers tuned for narrow individual bands
- Average receiver noise temperatures of 7 K--8 K below 18 GHz, 10 K--20 K from 18 GHz--35 GHz. ***For new Wideband LNA, aim is no more than 20% higher.***
- Goal is ~8 K--24 K from 8 GHz--35 GHz, in one band (not individual receivers)



# Cryogenic Wideband LNA Development

## Objectives and Challenges

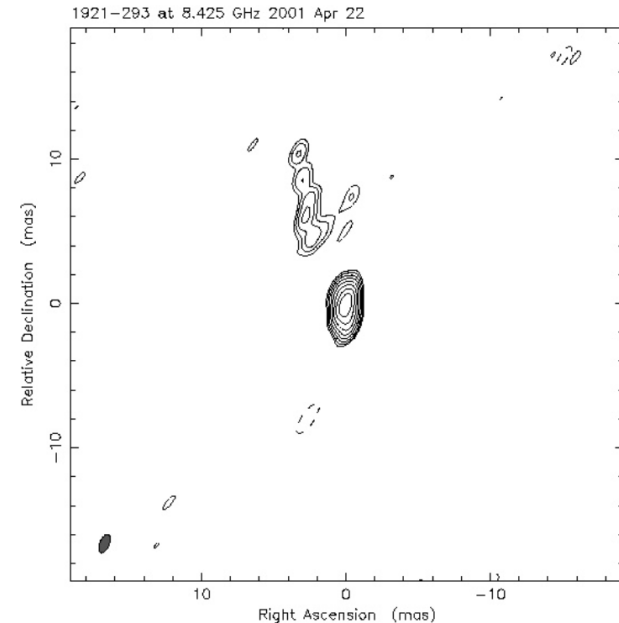
- Challenging to achieve low noise LNAs over 8 GHz--35 GHz
- JPL has LNA chips from multiple projects, some may be suitable for 8-35 GHz and will be tested
- *JPL is world leading in cryogenic MMIC design from 40 GHz--300 GHz*



Measured results from 8 GHz--50 GHz LNA from JPL (Bowen 2018) →  
*prove concept is possible*

# Wide-Band VLBA Receiver System for Astrometry and Navigation

- **Imaging:**  $\Delta$ DOR navigations benefit from *source images* in order to select most compact components
- **“Structure” effects**  $\ll \Delta$ DOR 1 nanoradian level
- DSN has only single baselines which prevents source imaging
- At present, **no** instrument in the world that can image at VLBI resolutions *in Ka band*



- **Ephemeris development:** Phase referencing at X- and Ka bands with the VLBA of spacecraft in planetary orbit relative to quasar can track planet's orbit with spacecraft without  $\Delta$ DOR tones and ***without disrupting telemetry***

# Wide-Band VLBA Receiver System for Astrometry and Navigation

Schedule: Year 1

Milestones	Planned Date
Evaluate LNAs 8-35 GHz	Oct -Jan 2019
Design new optimized LNAs & fab	Feb - June 2019
Design Receiver system compatible VLBA	Mar-Jun 2019
Design Wideband Feed compatible - VLBA	July 2019
Fabricate Feed	Aug 2019
Build Receiver	Aug 2019
Test new LNAs	Aug 2019
Test full receiver	Sep 2019



# Summary

## Wide-Band VLBA Receiver System for Astrometry and Navigation

- Prototype receiver system to provide reference frames for navigation and astrometry
- Critical components
  - Wideband LNA from 8 GHz--35 GHz
  - Wideband dual polarization feed compatible with VLBA

